The questions raised by cosmology are difficult and profound, and may seem removed from everyday "reality." We can always say, "the Sun is shining, it's going to burn on for an unimaginably long time, all is well." Nonetheless, the questions of cosmology are deep ones that fascinate the human intellect. One aspect that is especially intriguing is this: calculations on the formation and evolution of the universe have been performed that deliberately varied the values—just slightly—of certain fundamental physical constants. The result? A universe in which life as we know it could not exist. [For example, if the difference in mass between proton and neutron were zero, or small (less than the mass of the electron,  $0.511 \, \mathrm{MeV}/c^2$ ), there would be no atoms: electrons would be captured by protons never to be freed again.] Such results have given rise to a philosophical idea called the **Anthropic principle**, which says that if the universe were even a little different than it is, we could not be here. It might even seem that the universe is exquisitely tuned, almost as if to accommodate us.

Anthropic principle

## Summary

The night sky contains myriads of stars including those in the Milky Way, which is a "side view" of our **Galaxy** looking along the plane of the disc. Our Galaxy includes about 10<sup>11</sup> stars. Beyond our Galaxy are billions of other galaxies.

Astronomical distances are measured in **light-years** (1 ly  $\approx 10^{13}$  km). The nearest star is about 4 ly away and the nearest large galaxy is 2 million ly away. Our Galactic disc has a diameter of about 100,000 ly. Distances are often specified in **parsecs**, where 1 parsec = 3.26 ly.

Stars are believed to begin life as collapsing masses of hydrogen gas (protostars). As they contract, they heat up (potential energy is transformed to kinetic energy). When the temperature reaches about 10 million degrees, nuclear fusion begins and forms heavier elements (nucleosynthesis), mainly helium at first. The energy released during these reactions heats the gas so its outward pressure balances the inward gravitational force, and the young star stabilizes as a main-sequence star. The tremendous luminosity of stars comes from the energy released during these thermonuclear reactions. After billions of years, as helium is collected in the core and hydrogen is used up, the core contracts and heats further. The envelope expands and cools, and the star becomes a red giant (larger diameter, redder color). The next stage of stellar evolution depends on the mass of the star, which may have lost much of its original mass as its outer envelope escaped into space. Stars of residual mass less than about 1.4 solar masses cool further and become white dwarfs, eventually fading and going out altogether. Heavier stars contract further due to their greater gravity: the density approaches nuclear density, the huge pressure forces electrons to combine with protons to form neutrons, and the star becomes essentially a huge nucleus of neutrons. This is a neutron star, and the energy released from its final core collapse is believed to produce supernovae explosions. If the star is very massive, it may contract even further and form a black hole, which is so dense that no matter or light can escape from it.

In the general theory of relativity, the equivalence principle states that an observer cannot distinguish acceleration from a gravitational field. Said another way, gravitational and inertial masses are the same. The theory predicts gravitational bending of light rays to a degree consistent with experiment. Gravity is treated as a curvature in space and time, the curvature being greater near massive bodies. The universe as a whole may be curved. With sufficient mass, the curvature of the universe would be positive, and the universe is *closed* and *finite*; otherwise, it would be *open* and *infinite*.

Distant galaxies display a **redshift** in their spectral lines, interpreted as a Doppler shift. The universe seems to be **expanding**, its galaxies racing away from each other at speeds (v) proportional to the distance (d) between them:

$$v = Hd, (33-6)$$

which is known as **Hubble's law** (*H* is the **Hubble parameter**). This expansion of the universe suggests an explosive origin, the **Big Bang**, which occurred about 13.7 billion years ago.

The **cosmological principle** assumes that the universe, on a large scale, is homogeneous and isotropic.

Important evidence for the Big Bang model of the universe was the discovery of the **cosmic microwave background** radiation (CMB), which conforms to a blackbody radiation curve at a temperature of 2.725 K.

The standard model of the Big Bang provides a possible scenario as to how the universe developed as it expanded and cooled after the Big Bang. Starting at 10-43 seconds after the Big Bang, according to this model, there was a series of phase transitions during which previously unified forces of nature "condensed out" one by one. The inflationary scenario assumes that during one of these phase transitions, the universe underwent a brief but rapid exponential expansion. Until about 10<sup>-35</sup> s, there was no distinction between quarks and leptons. Shortly thereafter, quarks were confined into hadrons (the hadron era). About 10<sup>-4</sup> s after the Big Bang, the majority of hadrons disappeared, having combined with anti-hadrons, producing photons, leptons and energy, leaving mainly photons and leptons to freely move, thus introducing the lepton era. By the time the universe was about 10 s old, the electrons too had mostly disappeared, having combined with their antiparticles; the universe was radiation-dominated. A couple of minutes later, nucleosynthesis began, but lasted only a few minutes.