the projections at angle 2. We calculate the difference between the measured projections at angle 2 (6 and 14) and the projections based on the previous estimate (top row: $3\frac{1}{2} + 6\frac{1}{2} = 10$; same for bottom row). Then we distribute this difference equally to the squares in that row. For the top row, we have

$$3\frac{1}{2} + \frac{6-10}{2} = 1\frac{1}{2}$$
 and $6\frac{1}{2} + \frac{6-10}{2} = 4\frac{1}{2}$;

and for the bottom row,

$$3\frac{1}{2} + \frac{14 - 10}{2} = 5\frac{1}{2}$$
 and $6\frac{1}{2} + \frac{14 - 10}{2} = 8\frac{1}{2}$

These values are inserted as shown in Fig. 25-46c. Next, the projection at angle 3 gives

(upper left)
$$1\frac{1}{2} + \frac{11-10}{2} = 2$$
 and (lower right) $8\frac{1}{2} + \frac{11-10}{2} = 9$;

and that for angle 4 gives

(lower left)
$$5\frac{1}{2} + \frac{9-10}{2} = 5$$
 and (upper right) $4\frac{1}{2} + \frac{9-10}{2} = 4$.

The result, shown in Fig. 25–46d, corresponds exactly to the true values. (In real situations, the true values are not known, which is why these computer techniques are required.) To obtain these numbers exactly, we used six pieces of information (two each at angles 1 and 2, one each at angles 3 and 4). For the much larger number of pixels used for actual images, exact values are generally not attained. Many iterations may be needed, and the calculation is considered sufficiently precise when the difference between calculated and measured projections is sufficiently small. The above example illustrates the "convergence" of the process: the first iteration (b to c in Fig. 25–46) changed the values by 2, the last iteration (c to d) by only $\frac{1}{2}$.

Summary

A **camera** lens forms an image on film, or on a charge-coupled device in a digital camera, by allowing light in through a shutter. The lens is focused by moving it relative to the film, and its f-stop (or lens opening) must be adjusted for the brightness of the scene and the chosen shutter speed. The f-stop is defined as the ratio of the focal length to the diameter of the lens opening.

The human **eye** also adjusts for the available light—by opening and closing the iris. It focuses not by moving the lens, but by adjusting the shape of the lens to vary its focal length. The image is formed on the retina, which contains an array of receptors known as rods and cones.

Diverging eye-glass or contact lenses are used to correct the defect of a nearsighted eye, which cannot focus well on distant objects. Converging lenses are used to correct for defects in which the eye cannot focus on close objects.

A simple magnifier is a converging lens that forms a virtual image of an object placed at (or within) the focal point. The angular magnification, when viewed by a relaxed normal eye, is

$$M = \frac{N}{f}, (25-2a)$$

where f is the focal length of the lens and N is the near point of the eye (25 cm for a "normal" eye).

An **astronomical telescope** consists of an **objective lens** or mirror, and an **eyepiece** that magnifies the real image formed by the objective. The **magnification** is equal to the ratio of the objective and eyepiece focal lengths, and the image is inverted:

$$M = -\frac{f_0}{f_c}. (25-3)$$

[*A compound **microscope** also uses objective and eyepiece lenses, and the final image is inverted. The total magnification is the product of the magnifications of the two lenses and is approximately

$$M \approx \frac{Nl}{f_c f_o}$$
, (25-6b)

where l is the distance between the lenses, N is the near point of the eye, and $f_{\rm o}$ and $f_{\rm e}$ are the focal lengths of objective and eyepiece, respectively.]

Microscopes, telescopes, and other optical instruments are limited in the formation of sharp images by lens aberrations. These include spherical aberration, in which rays passing through the edge of a lens are not focused at the same point as those that pass near the center; and chromatic aberration, in which different colors are focused at different points. Compound lenses, consisting of several elements, can largely correct for aberrations.

The wave nature of light also limits the sharpness, or **resolution**, of images. Because of diffraction, it is *not possible* to discern details smaller than the wavelength of the radiation being used. This limits the useful magnification of a light microscope to about 500×.

[*X-rays are a form of electromagnetic radiation of very short wavelength. They are produced when high-speed electrons, accelerated by high voltage in an evacuated tube, strike a glass or metal target.]

[*Computed tomography (CT or CAT scans) uses many narrow X-ray beams through a section of the body to construct an image of that section.]